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# New stationary solutions of the nonlinear drift wave equation

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#### Abstract:

A new type of nonlinear wave modes which occur as the electrostatic drift waves in an inhomogeneous magnetized plasma is presented. The existence of a new type of spiky solitary wave and an explosive mode with the negative potential are predicted as stationary solutions of this equation. These solutions are a consequence of a density gradient and not connected with a temperature gradient. Using these nonlinear wave modes, the solitary structure and the explosive event concerning nonlinear drift waves propagating in interplanetary space and the Earth's magnetosphere, are understandable.

Key words: nonlinear drift wlave equation, new type of nonlinear wave modes, spiky solitary wave, explosive mode

#### 1. Introduction

Drift waves in the inhomogeneous magnetized plasmas have been studied experimentally and theoretically in the context of vortexes and The drift solitary waves have been experimentally solitary waves. observed in the Q-machine [1]. Simple two-dimensional drift wave turbulence models based on wave-wave nonlinear coupling have been examined with numerical techniques [2]. Recently, Lakhin et al. theoretically studied the low frequency drift solitary wave [3]. In the situation where the nonlinear drift wave propagates in interplanetary space, it may be possible that spiky solitary waves and explosive (bursting) events are detected by satellites [4]. Hendel et al. suggest that mode stabilization occurs because the mode density pertur bations affect the diamagnetic drift [1]. Hence, it is possible for new type of nonlinear wave modes to occur in drift waves. stresses the importance of the higher-order nonlinearity with regard to the collapse of Langmuir waves, and mentions that the waves with ne gative energy have the explosive instability [5]. On the other hand, several studies of the nonlinear evolution equation have been made in the context of nonlinear plasma waves [6-8]. Su et al. show plasma waves associated with the exponential nonlinearity [9-10]. However. the new type of nonlinear drift waves have not yet been shown.

In this paper, we consider a nonlinear drift wave equation in an inhomogenious magnetized plasma as a model equation, that is, Hasegawa-Mima equation. The purpose of this investigation is to demonstrate a new type of spiky solitary wave and an explosive mode of the nonlinear drift wave equation. It will be expected that these solutions extend the scheme of well established nonlinear evolution equations and this investigation may be applied to explaining the behaviour of larger

amplitude nonlinear waves propagating in plasmas. We do not compare the present results to collisional plasmas.

The layout of this paper is as follows. In section 2, we consider the stationary nonlinear drift wave equation propagating in an inhomogeneous magnetized plasma. A nonlinear evolution equation is derived by several approximations. In section 3, new stationary wave solutions of this equation are shown. These are a spiky solitary wave and an explosive mode. The last section is devoted to concluding discussion.

#### 2. Derivation of a nonlinear evolution equation

We begin by considering a nonlinear drift(Hasegawa-Mima) equation [11,12] as a model:

$$\frac{\partial}{\partial t} (\nabla^2 \phi - \phi) - (\nabla \phi \times \overrightarrow{i}_z) \cdot \nabla (\nabla^2 \phi - \log n_0) = 0 \qquad . \tag{1}$$

We take the following assumptions:

- 1. the magnetic field  $B = B_0 i_z$  is assumed to be constant and homogeneous.
- 2. the density  $n_0$  is inhomogeneous in the x-direction and depends strongly on the spatial variation.
- 3. the drift wave is not connected with the temperature gradient.
- 4. the nonlinearity competes with the dispersion effect.

When the density  $n_0 = n_0(x)$  varying in a direction perpendicular

to the magnetic field  $B=B_0$   $i_z$  in the z-direction, eq.(1) is reduced to

$$\frac{\partial}{\partial t} \left( \nabla^2 \phi - \phi - \log n_0 \right) + \left[ \phi, \nabla^2 \phi - \phi - \log n_0 \right] = 0$$
 (2)

where

$$[\alpha,\beta] = \frac{\partial \alpha}{\partial x} \frac{\partial \beta}{\partial y} - \frac{\partial \alpha}{\partial y} \frac{\partial \beta}{\partial x}.$$

In order to obtain stationary wave solutions, we put

$$\phi = \phi (x, \eta)$$

$$\eta = y - u t \qquad . \tag{3}$$

Then eq. (2) becomes

$$[\phi - u x, \nabla^2 \phi - \phi - \log n_0] = 0 \qquad . \tag{4}$$

Equation (4) has the integral

$$\nabla^2 \phi - \phi - \log n_0 = F(\phi - u x) \tag{5}$$

where  $F(\phi-u\,x)$  is an arbitrary function of the argument. Such solutions travel with the velocity u in the y direction. In the case of solitary localized solutions,  $F(\phi-u\,x)$  is determined from the condition that  $\phi\to 0$  at  $|y|\to \infty$  for the fixed values of x.

Then we obtain, from eq. (5),

$$\nabla^2 \phi - \phi + \log \left[ n_0 \left[ x - \frac{\phi}{u} \right] \right] - \left[ \log n_0(x) \right] = 0$$
 (6)

We expand  $\log (n_0(x-\phi/u))$  in the power series of  $\phi$  as follows

$$\log\left[n_0\left(x-\frac{\phi}{u}\right)\right]-\log n_0(x)$$

$$= -\frac{\phi}{u} \left[ \frac{\partial}{\partial x} \log n_{\circ} \right]_{x=0} - x \frac{\phi}{u} \left[ \frac{\partial^{2}}{\partial x^{2}} \log n_{\circ} \right]_{x=0}$$

$$+ \frac{\phi^{2}}{2 u^{2}} \left[ \frac{\partial^{2}}{\partial x^{2}} \log n_{\circ} \right]_{x=0}$$

$$+ \left[ -\frac{1}{2} x^{2} \frac{\phi}{u} + \frac{1}{2} x \frac{\phi^{2}}{u^{2}} - \frac{1}{6} \frac{\phi^{3}}{u^{3}} \right] \left[ \frac{\partial^{3}}{\partial x^{3}} \log n_{\circ} \right]_{x=0}$$

$$+ \left[ -\frac{1}{6} x^{3} \frac{\phi}{u} + \frac{1}{4} x^{2} \frac{\phi^{2}}{u^{2}} - \frac{1}{6} x \frac{\phi^{3}}{u^{3}} + \frac{1}{24} \frac{\phi^{4}}{u^{4}} \right]$$

$$\times \left[ \frac{\partial^{4}}{\partial x^{4}} \log n_{\circ} \right]_{x=0} + \cdots$$

$$= -\left\{ \left[ \frac{\partial}{\partial x} \log n_{\circ} \right]_{x=0} + x \left[ \frac{\partial^{2}}{\partial x^{2}} \log n_{\circ} \right]_{x=0} + \cdots \right\}$$

$$+ \frac{1}{2} x^{2} \left[ \frac{\partial^{3}}{\partial x^{3}} \log n_{\circ} \right]_{x=0} + x \left[ \frac{\partial^{3}}{\partial x^{3}} \log n_{\circ} \right]_{x=0} \right\}$$

$$+ \frac{1}{2} \left\{ \left[ \frac{\partial^{2}}{\partial x^{2}} \log n_{\circ} \right]_{x=0} + x \left[ \frac{\partial^{3}}{\partial x^{3}} \log n_{\circ} \right]_{x=0} \right\} \frac{\phi^{2}}{u^{2}}$$

$$- \frac{1}{6} \left\{ \left[ \frac{\partial^{3}}{\partial x^{3}} \log n_{\circ} \right]_{x=0} + x \left[ \frac{\partial^{4}}{\partial x^{4}} \log n_{\circ} \right]_{x=0} \right\} \frac{\phi^{2}}{u^{3}}$$

$$+ \frac{1}{24} \left\{ \left[ \frac{\partial^{4}}{\partial x^{4}} \log n_{\circ} \right]_{x=0} \right\} \frac{\phi^{4}}{u^{4}}$$

$$+ O(\phi^{5}) + \cdots$$
(7)

Substituting eq. (7) into eq. (6), one obtain

 $\nabla^2 \phi$ 

$$= \left[1 + \frac{1}{u} \left\{ \left[\frac{\partial}{\partial x} \log n_{0}\right]_{x=0} + x \left[\frac{\partial^{2}}{\partial x^{2}} \log n_{0}\right]_{x=0} + \frac{1}{2} x^{2} \left[\frac{\partial^{3}}{\partial x^{3}} \log n_{0}\right]_{x=0} + \frac{1}{6} x^{3} \left[\frac{\partial^{4}}{\partial x^{4}} \log n_{0}\right]_{x=0} \right\} \right] \phi$$

$$- \frac{1}{2} \frac{1}{u^{2}} \left\{ \left[\frac{\partial^{2}}{\partial x^{2}} \log n_{0}\right]_{x=0} + x \left[\frac{\partial^{3}}{\partial x^{3}} \log n_{0}\right]_{x=0} + \frac{1}{2} x^{2} \left[\frac{\partial^{4}}{\partial x^{4}} \log n_{0}\right]_{x=0} \right\} \phi^{2}$$

$$+ \frac{1}{6} \frac{1}{u^{3}} \left\{ \left[\frac{\partial^{3}}{\partial x^{3}} \log n_{0}\right]_{x=0} + x \left[\frac{\partial^{4}}{\partial x^{4}} \log n_{0}\right]_{x=0} \right\} \phi^{3}$$

$$- \frac{1}{24} \frac{1}{u^{4}} \left\{ \left[\frac{\partial^{4}}{\partial x^{4}} \log n_{0}\right]_{x=0} \right\} \phi^{4} .$$

$$+ O(\phi^{5}) + \cdots$$

$$(8)$$

Since the behaviour of stationary wave solutions is determined largely by the spatial dependence of the density  $n_0 = n_0(x)$ , we assume a simple exponential density profile,

$$n_0(x) \sim \exp(-\varepsilon_n x)$$
.

and constant temperature. Here  $\varepsilon_n$  is the usual drift theory expansion parameter  $\varepsilon_n = \rho_0/L_n$ , where  $\rho_0 = c_s/\omega_1 = (\kappa T_e/m_i)^{1/2}/(eB/m_1)$  is the Larmor radius, and  $L_n$  is the density gradient scale length.

As is seen in the observation,  $c_{\rm s}$  and  $\omega_{\rm i}$  are the sound velocity and the ion cyclotron frequency.

Using this model, we can describe eq.(8) to

$$\nabla^2 \phi = c_1(x) \phi + c_2(x) \phi^2 + c_3(x) \phi^3 + c_4(x) \phi^4 + O(\phi^5) + \cdots,$$
(9)

where the coefficients are determined to

$$c_{1}(x)=1+\frac{1}{u}\left\{\left[\frac{\partial}{\partial x}\log n_{0}\right]_{x=0}+x\left[\frac{\partial^{2}}{\partial x^{2}}\log n_{0}\right]_{x=0}+\frac{1}{2}x^{2}\left[\frac{\partial^{3}}{\partial x^{3}}\log n_{0}\right]_{x=0}+\frac{1}{6}x^{3}\left[\frac{\partial^{4}}{\partial x^{4}}\log n_{0}\right]_{x=0}\right\},$$

$$c_{2}(x)=-\frac{1}{2u^{2}}\left\{\left[\frac{\partial^{2}}{\partial x^{2}}\log n_{0}\right]_{x=0}+x\left[\frac{\partial^{3}}{\partial x^{3}}\log n_{0}\right]_{x=0}+\frac{1}{2}x^{2}\left[\frac{\partial^{4}}{\partial x^{4}}\log n_{0}\right]_{x=0}\right\},$$

$$c_{3}(x)=\frac{1}{6u^{3}}\left\{\left[\frac{\partial^{3}}{\partial x^{3}}\log n_{0}\right]_{x=0}+x\left[\frac{\partial^{4}}{\partial x^{4}}\log n_{0}\right]_{x=0}\right\},$$

$$c_{4}(x)=-\frac{1}{2u^{4}}\left\{\left[\frac{\partial^{4}}{\partial x^{4}}\log n_{0}\right]_{x=0}\right\},$$

where  $c_1(x)>0$ ,  $c_2(x)<0$ ,  $c_3(x)>0$  and  $c_4(x)<0$ .

Since eq. (9) heavily depends on x, we can approximate

$$\left| \begin{array}{c} \partial \phi \\ \overline{\partial \eta} \end{array} \right| \; \ll \; \left| \begin{array}{c} \partial \phi \\ \overline{\partial x} \end{array} \right| \;\; .$$

Then the left hand-side of eq.(9) is reduced to  $\nabla^2 \phi \simeq (\partial^2 \phi / \partial x^2)$ . This means that we consider the quasi-one dimensional case. Using these approximations and integrating eq.(9), we obtain

$$\left(\begin{array}{c} \frac{\partial \phi}{\partial x} \end{array}\right)^{2} = c_{1}(x)\phi^{2} + \frac{2}{3}c_{2}(x)\phi^{3} + \frac{1}{2}c_{3}(x)\phi^{4} + \frac{2}{5}c_{4}(x)\phi^{5}$$

$$\equiv -\mathcal{U}(\phi)$$

$$= \phi^{2}(a_{1} - a_{2}\phi - a_{3}\phi^{2} - a_{4}\phi^{3}) \tag{10}$$

where

$$a_1 = c_1(x)$$
,  
 $a_2 = -\frac{2}{3}c_2(x)$ ,  
 $a_3 = -\frac{1}{2}c_3(x)$ ,  
 $a_4 = -\frac{2}{5}c_4(x)$ .

If  $a_4\phi^5\approx 0$  in eq.(10), the solution is already shown. However, the solution of this equation which includes the nonlinear term  $a_4\phi^5$  has not yet been solved.

#### 3. Stationary solutions of eq. (10)

In order to obtain stationary wave solutions of eq. (10), we transform eq. (10) to

$$\left(\begin{array}{c} \frac{\partial \phi}{\partial x} \end{array}\right)^{2} \equiv -\mathcal{U}(\phi)$$

$$= A \phi^{2} (\varphi_{0} - \phi)^{3} , \qquad (11)$$

because the solutions of eq.(10) exist only in the region of  $-\mathcal{U}(\phi) > 0$ , where we determine that

$$A = a_4 \quad , \qquad \qquad \varphi_0 = -\frac{a_3}{3 a_4} \quad .$$

A relationship between the potential function  $-\mathcal{U}(\phi)$  and  $\phi$  is shown in Fig.1. One can understand that the proper localized solutions are obtained in the upper region of the horizontal axis. According to eq. (11),  $d\phi/dx$  is finite everywhere  $\phi$  is finite. It should be noted that the potential function  $\mathcal{U}(\phi)$  is finite in the region  $\phi>0$ , but is infinite in the region  $\phi<0$ . As a result of this, we have to integrate eq.(10) independently in the two regions.

Imposing the boundary conditions  $\phi$ ,  $d\phi/dx$ ,  $d^2\phi/dx^2 \rightarrow 0$  at  $|x| \rightarrow \infty$ , we have, from eq.(11),

$$\pm A^{1/2}(x-x_0) = \int \frac{d\phi}{\phi (\varphi_0 - \phi)^{3/2}}$$

$$= \begin{cases} \int_{\phi}^{\phi_0} \dots & \text{in the region } 0 < \phi < \varphi_0 \\ \int_{\phi}^{0} \dots & \text{in the region } -\infty < \phi < 0 \end{cases}$$
 (12)

where  $\phi_0$  denotes the maximum amplitude of the drift wave at  $x \to x_0$ .

We should note the sign of the coefficients of eq.(10). If the coefficients  $a_1>0$ ,  $a_2>0$ ,  $a_3<0$ ,  $a_4>0$ , then u>0 and  $\varphi_0>0$ .

Integrating eq. (12) and using the boundary conditions, we have

$$\pm \frac{A^{1/2}}{2} \varphi_0^{3/2} (x - x_0)$$

$$= -\left[\frac{\varphi_0}{\varphi_0 - \phi_s(\varphi_0, \phi_0, x)}\right]^{1/2} + \operatorname{arctanh}\left[\frac{\varphi_0 - \phi_s(\varphi_0, \phi_0, x)}{\varphi_0}\right]^{1/2} + \mathcal{H}_1(\varphi_0, \phi_0), \qquad (14)$$

where

$$\mathcal{H}_{1}(\varphi_{0}, \phi_{0}) = \left[\frac{\varphi_{0}}{\varphi_{0} - \phi_{0}}\right]^{1/2} - \operatorname{arctanh}\left[\frac{\varphi_{0} - \phi_{0}}{\varphi_{0}}\right]^{1/2}.$$

We obtain the solution of eq. (12) as follows

$$\phi_s(\varphi_0,\phi_0,x)$$

$$= \varphi_{0} \operatorname{sech}^{2} \left[ \left[ \frac{\varphi_{0}}{\varphi_{0} - \varphi_{s}(\varphi_{0}, \varphi_{0}, x)} \right]^{1/2} \pm \frac{A^{1/2}}{2} \varphi_{0}^{3/2}(x - x_{0}) - \mathcal{H}_{1}(\varphi_{0}, \varphi_{0}) \right]$$
(15)

Since  $\mathcal{H}_1(\varphi_0, \phi_0)$  of eq.(15) connects the wave profile in x>0 with that in x<0, the wave profile of  $\phi_s(\varphi_0, \phi_0, x)$  is continuous and finite at  $x\to x_0$ . Equation (15) forms a new type of spiky solitary wave.  $\phi_s(\varphi_0, \phi_0, x)$  is the amplitude of the spiky solitary wave. We illustrate the wave profile of eq.(15) in Fig.2.

Similarly, calculating eq. (13), we get

$$\pm \frac{A^{1/2}}{2} \varphi_0^{3/2} (x - x_0)$$

$$= -\left[\frac{\varphi_0}{\varphi_0 - \phi_E(\varphi_0, \phi_0, x)}\right]^{1/2} + \operatorname{arccoth}\left[\frac{\varphi_0 - \phi_E(\varphi_0, \phi_0, x)}{\varphi_0}\right]^{1/2}$$

$$+ \mathcal{H}_2(\varphi_0, \phi_0), \qquad (16)$$

where

$$\mathscr{H}_{2}(\varphi_{0}, \phi_{0}) = \left[\frac{\varphi_{0}}{\varphi_{0} - \phi_{0}}\right]^{1/2} - \operatorname{arccoth}\left[\frac{\varphi_{0} - \phi_{0}}{\varphi_{0}}\right]^{1/2}.$$

We then obtain the solution of eq. (13), from eq. (16), as

$$\phi_{\rm E}(\varphi_{\rm o},\phi_{\rm o},x)$$

The potential function  $\mathcal{U}(\phi)$  is infinite in the region  $\phi < 0$  in Fig.1. Hence, it is predicted that eq.(13) has no finite solutions. Calculating eq.(17) numerically, we can obtain the solution that the amplitude becomes infinite.  $\phi_{\rm E}(\varphi_{\rm O},\phi_{\rm O},x)$  denotes the amplitude of the explosive solution. Equation (17) takes an explosive profile at  $\pm (1/2) A^{1/2} \varphi_{\rm O}^{3/2} (x-x_{\rm O}) \rightarrow \mathcal{H}_2(\varphi_{\rm O},\phi_{\rm O})$ , which is illustrated in Fig.3.

#### 4. Concluding discussion

We present new type of nonlinear wave modes associated with in nonlinear drift waves. We show that a new nonlinear evolution equation is derived from nonlinear drift wave equation, assuming the spatial dependence of the density and using a quasi one-dimensional approxi-This equation has the fifth-order nonlinear potential term. The behaviour of the plasma physical system in which the fifth-order nonlinearity competes with the dispersion effect is described by eq. (10).The stationary wave solutions of eq. (10) with the nonlinear potential term  $\phi^{5}$  is shown for the first time in this investigation. The solution of eq. (11) is integrable according to the selection of the proper boundary conditions and the proper integration region. stationary wave solutions of this equation are obtained due to fifth-order nonlinearity, and bear a spiky solitary wave solution and an explosive solution. These solutions are first shown in drift It should be noted that the explosive solution eventually will make the ordering in the derivation of the higher-order nonlinear evolution equation break down. In addition, the explosive solution is associated with the wave with negative potential. It is worth noticing that the explosive mode presented here is similar to derived in the potential structure of Zakharov.

In actual situations, spiky solitary waves and solar radio burst events are frequently observed in interplanetary space. Hence, referring to the present spiky solitary waves and explosive(bursting) modes, we can understanding the properties concerning the new nonlinear wave modes of nonlinear drift waves in space plasmas. Although the author has not examined the application of these results to a specific observational result, this investigation is important in discussing

the nonlinear wave modes which occur in plasmas. This theory is therefore applicable to another nonlinear waves in physical systems.

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## Captions of figures

- Fig.1 The potential function  $-\mathscr{U}$  versus potential  $\phi$ .
- Fig. 2 The profile of a spiky solitary wave described by eq.(15).
- Fig. 3 The profile of an explosive solution described by eq.(17).

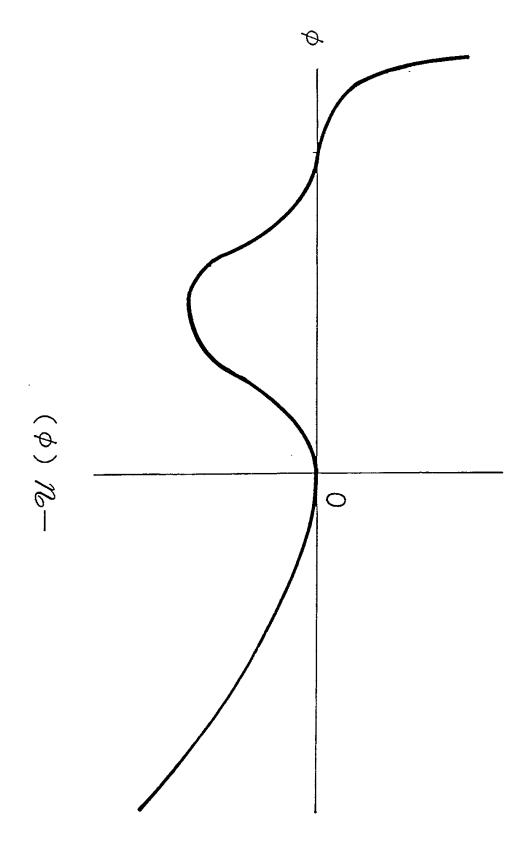


Fig.1

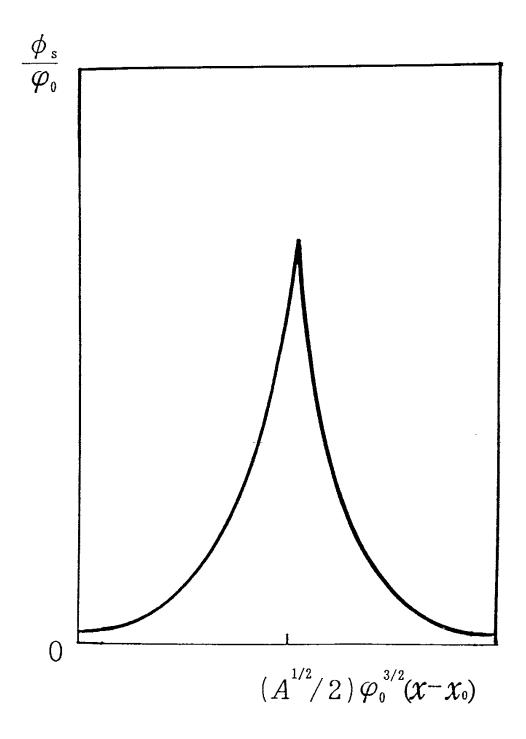


Fig.2

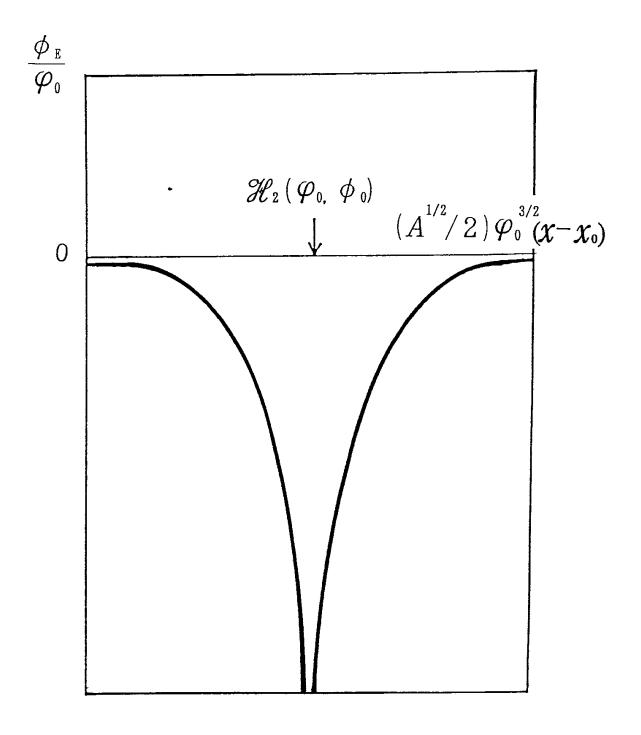


Fig.3

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